

# Relaxation techniques for the Euler equations with gravity

**Vivien Desveaux**

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Joint work with:

- Markus Zenk (Würzburg)
- Christian Klingenberg (Würzburg)
- Christophe Berthon (Nantes)

# Outline

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# The system of Euler equation with gravity

$$\begin{cases} \partial_t \rho + \partial_x \rho u = 0 \\ \partial_t \rho u + \partial_x (\rho u^2 + p) = -\rho \partial_x \Phi \\ \partial_t E + \partial_x (E + p)u = -\rho u \partial_x \Phi \end{cases}$$

- $\rho$  : density  
 $u$  : velocity  
 $E = \rho e + \rho u^2/2$  : total energy, with  $e$  the internal energy  
 $\tau := 1/\rho$ : specific volume  
 $p = p(\tau, e)$  : pressure given by a general law  
 $\Phi(x)$  : given smooth gravitational potential
- Hyperbolicity assumption:

$$c^2 := \tau^2 (p \partial_e p - \partial_\tau p) > 0$$

# The system of Euler equation with gravity

$$\begin{cases} \partial_t \rho + \partial_x \rho u = 0 \\ \partial_t \rho u + \partial_x (\rho u^2 + p) = -\rho \partial_x \Phi \\ \partial_t E + \partial_x (E + p)u = -\rho u \partial_x \Phi \end{cases}$$

- We define

- ▶ the vector of conservative variables  $w = (\rho, \rho u, E)^T$ ,
- ▶ the flux function  $f(w) = (\rho u, \rho u^2 + p, (E + p)u)^T$ ,
- ▶ the source term  $s(w) = (0, -\rho, -\rho u)^T$ ,

to rewrite the system into the compact form

$$\partial_t w + \partial_x f(w) = s(w) \partial_x \Phi.$$

- The set of physical admissible states is

$$\Omega = \left\{ w \in \mathbb{R}^3, \quad \rho > 0, \quad e > 0 \right\}.$$

# Steady states

The steady states at rest are described by

$$\begin{cases} u = 0 \\ \partial_x p = -\rho \partial_x \Phi \end{cases}$$

No explicit expression of all the steady states at rest

## Hydrostatic atmosphere

$$\begin{cases} u(x) = 0, \\ \rho(x) = \alpha e^{-\beta \Phi(x)}, \\ p(x) = \frac{\alpha}{\beta} e^{-\beta \Phi(x)}, \end{cases} \quad \text{with } \alpha > 0 \text{ and } \beta > 0$$

# Entropy

The pressure is assumed to satisfy the second law of thermodynamics  
 $\Rightarrow$  existence of a specific entropy  $s(\tau, e)$  which satisfies

$$-Tds = de + pd\tau, \quad \text{with } T > 0 \text{ the temperature}$$

## Lemma

*The smooth solutions satisfy the additional conservation laws*

$$\partial_t \rho \mathcal{F}(s) + \partial_x \rho \mathcal{F}(s) u = 0,$$

*for all smooth function  $\mathcal{F}$ .*

*Moreover the application  $w \mapsto \rho \mathcal{F}(s)$  is strictly convex iff*

$$\mathcal{F}'(s) > 0 \quad \text{and} \quad \frac{1}{c_p} \mathcal{F}'(s) + \mathcal{F}''(s) > 0.$$

# Objectives

Derive a numerical scheme which has the following properties:

- Preservation of the set  $\Omega$
- Accurate approximation of all the steady states at rest
- Exact capture of the hydrostatic atmosphere
- Discrete entropy inequalities
- General gravitational potential and general pressure law

## Means

- Finite volume method (approximate Riemann solver)
- Relaxation method

## References

- [Chalons et al. '10] : constant gravity field ( $\Phi(x) = gx$ )
- [Käpelli & Mishra '13] : isentropic steady states

# The relaxation method without source term

- Initial system:

$$\partial_t w + \partial_x f(w) = 0. \quad (1)$$

- Relaxation system:

$$\partial_t W + \partial_x F(W) = \frac{1}{\varepsilon} R(W), \quad (2)$$

- ▶ (2) should formally gives back (1) when  $\varepsilon \rightarrow 0$ .
  - ▶ (2) should be “simpler” than (1) (e.g. only linearly degenerate fields)

- The relaxation scheme is based on a splitting strategy:

**Time evolution:** We evolve the initial data by the Godunov scheme for the system  $\partial_t W + \partial_x F(W) = 0$  (i.e.  $\varepsilon = +\infty$ ).

$\Rightarrow$  We need the exact solution of the Riemann problem

**Relaxation:** We take into account the relaxation source term by solving  $\partial_t W = \frac{1}{\varepsilon} R(W)$  then taking the limit for  $\varepsilon \rightarrow 0$ .



# The Suliciu model([Suliciu '98], [Bouchut '04]...)

$$\left\{ \begin{array}{l} \partial_t \rho + \partial_x \rho u = 0 \\ \partial_t \rho u + \partial_x (\rho u^2 + \pi) = -\rho \partial_x \Phi \\ \partial_t E + \partial_x (E + \pi) u = -\rho u \partial_x \Phi \\ \partial_t \rho \pi + \partial_x (\rho \pi + a^2) u = \frac{\rho}{\varepsilon} (p(\tau, e) - \pi) \\ \partial_t \Phi = 0 \end{array} \right. \quad \begin{array}{l} \text{Equilibrium} \\ \\ \\ \pi = p(\tau, e) \end{array}$$

**Eigenvalues:**  $u - \frac{a}{\rho}, \quad 0, \quad u, \quad u + \frac{a}{\rho}$

**Advantage:**

- All the fields are linearly degenerate

**Difficulties:**

- The order of the eigenvalues is not fixed *a priori*
- The Riemann invariants for the eigenvalue 0 are strongly nonlinear

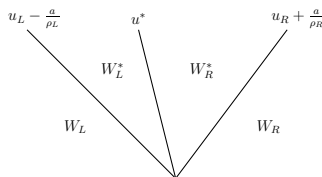
# Relaxation model with moving gravity

$$\left\{ \begin{array}{l} \partial_t \rho + \partial_x \rho u = 0 \\ \partial_t \rho u + \partial_x (\rho u^2 + \pi) = -\rho \partial_x Z \\ \partial_t E + \partial_x (E + \pi) u = -\rho u \partial_x Z \\ \partial_t \rho \pi + \partial_x (\rho \pi + a^2) u = \frac{\rho}{\varepsilon} (p(\tau, e) - \pi) \\ \partial_t \rho Z + \partial_x \rho Z u = \frac{\rho}{\varepsilon} (\Phi - Z) \end{array} \right. \quad \begin{array}{l} \text{Equilibrium} \\ \pi = p(\tau, e) \\ Z = \Phi \end{array}$$

- Eigenvalues:  $u - \frac{a}{\rho}$ ,  $u$ ,  $u + \frac{a}{\rho}$
- All the fields are linearly degenerate
- Fixed order of the eigenvalues:  $u - \frac{a}{\rho} < u < u + \frac{a}{\rho}$
- There is a **missing Riemann invariant** in order to determine a unique Riemann solution  
 $\Rightarrow$  We need a closure relation

# Relaxation model with moving gravity

$$\left\{ \begin{array}{l} \partial_t \rho + \partial_x \rho u = 0 \\ \partial_t \rho u + \partial_x (\rho u^2 + \pi) = -\rho \partial_x Z \\ \partial_t E + \partial_x (E + \pi) u = -\rho u \partial_x Z \\ \partial_t \rho \pi + \partial_x (\rho \pi + a^2) u = \frac{\rho}{\varepsilon} (p(\tau, e) - \pi) \\ \partial_t \rho Z + \partial_x \rho Z u = \frac{\rho}{\varepsilon} (\Phi - Z) \end{array} \right.$$



To approximate the equation

$$\partial_x p = -\rho \partial_x \Phi \quad \underset{\substack{\Longleftrightarrow \\ \text{equilibrium} \\ \pi = p(\tau, e) \\ Z = \Phi}}{\quad} \quad \partial_x \pi = -\rho \partial_x Z,$$

we propose the following closure relation:

$$\frac{\pi_R^* - \pi_L^*}{\Delta x} = -\bar{\rho}(\rho_L, \rho_R) \frac{Z_R - Z_L}{\Delta x},$$

where  $\bar{\rho}$  is a suitable  $\rho$ -average (defined later)

# Approximate Riemann solver

Equilibrium state:

$$W^{eq}(w) = (\rho, \rho u, E, \rho p(\tau, e), \rho \Phi)^T$$

## Theorem

*With the closure equation, the Riemann problem admits a unique solution  $W_{\mathcal{R}}\left(\frac{x}{t}, W_L, W_R\right)$ .*

*Moreover,*

$$w^{eq}\left(\frac{x}{t}, w_L, w_R\right) := W_{\mathcal{R}}^{(\rho, \rho u, E)}\left(\frac{x}{t}, W^{eq}(w_L), W^{eq}(w_R)\right)$$

*defines an approximate Riemann solver (in the sense of Harten, Lax and van Leer) for the Euler equations with gravity.*

# Complete relaxation reformulation

$$\left\{ \begin{array}{l} \partial_t \rho + \partial_x \rho u = 0 \\ \partial_t \rho u + \partial_x (\rho u^2 + \pi) = -\bar{\rho}(X^-, X^+) \partial_x Z \\ \partial_t E + \partial_x (E + \pi) u = -\bar{\rho}(X^-, X^+) u \partial_x Z \\ \partial_t \rho \pi + \partial_x (\rho \pi + a^2) u = \frac{\rho}{\varepsilon} (p(\tau, e) - \pi) \\ \partial_t \rho Z + \partial_x \rho Z u = \frac{\rho}{\varepsilon} (\Phi - Z) \\ \partial_t X^- + (u - \delta) \partial_x X^- = \frac{1}{\varepsilon} (\rho - X^-) \\ \partial_t X^+ + (u + \delta) \partial_x X^+ = \frac{1}{\varepsilon} (\rho - X^+) \end{array} \right.$$

Equilibrium

$$\pi = p(\tau, e) \quad Z = \Phi \quad X^\pm = \rho$$

- For  $\delta > 0$  small enough, the system is hyperbolic with eigenvalues

$$u - \frac{a}{\rho} < u - \delta < u < u + \delta < u + \frac{a}{\rho}.$$

- The system has a complete set of Riemann invariants.
- Leads to the same approximate Riemann solver  $w^{eq}(\frac{x}{t}, w_L, w_R)$ .

# Cargo-LeRoux formulation ( $\Phi(x) = gx$ )

We introduce a potential  $q = \int^x g\rho dx$ . Then we have

$$\partial_t q = \int^x g \partial_t \rho dx = - \int^x g \partial_x \rho u dx = -g\rho u = -u \partial_x q.$$

So  $q$  is governed by

$$\partial_t \rho q + \partial_x \rho q u = 0.$$

Equivalent reformulation of the Euler equations with gravity:

$$\begin{cases} \partial_t \rho + \partial_x \rho u = 0 \\ \partial_t \rho u + \partial_x (\rho u^2 + p) + \partial_x q = 0 \\ \partial_t E + \partial_x (E + p)u + u \partial_x q = 0 \\ \partial_t \rho q + \partial_x \rho q u = 0. \end{cases}$$

# Relaxation model with the Cargo-LeRoux formulation

- Case  $\Phi(x) = gx$  [Chalons et al. '10]

$$\begin{cases} \partial_t \rho + \partial_x \rho u = 0 \\ \partial_t \rho u + \partial_x (\rho u^2 + \pi) + \partial_x q = 0 \\ \partial_t E + \partial_x (E + \pi)u + u \partial_x q = 0 \\ \partial_t \rho q + \partial_x \rho q u = 0 \\ \partial_t \rho \pi + \partial_x (\rho \pi + a^2)u = \frac{\rho}{\varepsilon} (p(\tau, e) - \pi) \end{cases} \quad \begin{array}{l} \text{Equilibrium} \\ \\ \\ \\ \pi = p(\tau, e) \end{array}$$

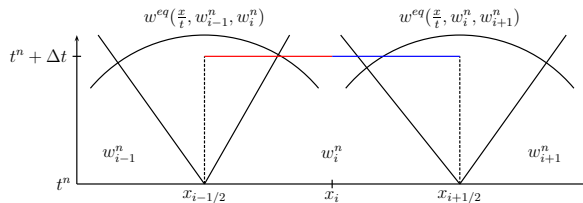
- Extension for a general gravitational field

If we define the potential by  $q = \int^x \rho \partial_x \Phi dx$ , it no longer satisfies a transport equation. We enforce the natural relaxation model

$$\begin{cases} \partial_t \rho + \partial_x \rho u = 0 \\ \partial_t \rho u + \partial_x (\rho u^2 + \pi) + \partial_x q = 0 \\ \partial_t E + \partial_x (E + \pi)u + u \partial_x q = 0 \\ \partial_t \rho q + \partial_x \rho q u = \frac{\rho}{\varepsilon} (\int^x \rho \partial_x \Phi dx - q) \\ \partial_t \rho \pi + \partial_x (\rho \pi + a^2)u = \frac{\rho}{\varepsilon} (p(\tau, e) - \pi) \end{cases} \quad \begin{array}{l} \text{Equilibrium} \\ \\ \\ q = \int^x \rho \partial_x \Phi dx \\ \pi = p(\tau, e) \end{array}$$

# The relaxation scheme

$w_i^n$ : approximation of the solution on the cell  $(x_{i-1/2}, x_{i+1/2})$  at time  $t^n$



CFL restriction

$$\frac{\Delta t}{\Delta x} \max_{i \in \mathbb{Z}} \left| u_i^n \pm \frac{a}{\rho_i^n} \right| \leq \frac{1}{2}$$

The update at time  $t^{n+1} = t^n + \Delta t$  is defined by

$$\begin{aligned} w_i^{n+1} = & \frac{1}{\Delta x} \int_{x_{i-1/2}}^{x_i} w^{eq} \left( \frac{x - x_{i-1/2}}{\Delta t}, w_{i-1}^n, w_i^n \right) dx \\ & + \frac{1}{\Delta x} \int_{x_i}^{x_{i+1/2}} w^{eq} \left( \frac{x - x_{i+1/2}}{\Delta t}, w_i^n, w_{i+1}^n \right) dx \end{aligned}$$



# Properties of the relaxation scheme (2)

## Theorem (Robustness)

Assume the parameter  $a$  satisfies the following inequalities:

$$u_L - \frac{a}{\rho_L} < u^* < u_R + \frac{a}{\rho_R},$$

$$e_L + \frac{\pi_L^{*2} - p_L^2}{2a^2} > 0, \quad e_R + \frac{\pi_R^{*2} - p_R^2}{2a^2} > 0$$

Then the relaxation scheme preserves the set of physical states  $\Omega$ .

## Theorem (Well-balancedness)

The relaxation scheme preserves exactly the initial data satisfying

$$\begin{cases} u_i^0 = 0, \\ \frac{p_{i+1}^0 - p_i^0}{\Delta x} = -\bar{\rho}(\rho_i^0, \rho_{i+1}^0) \frac{\Phi_{i+1} - \Phi_i}{\Delta x}. \end{cases}$$

Then  $w_i^{n+1} = w_i^n, \quad \forall i \in \mathbb{Z}, \quad \forall n \in \mathbb{N}$ .

# Properties of the relaxation scheme (1)

## Theorem (Exact preservation of the hydrostatic atmosphere)

Assume the average function  $\bar{\rho}$  is defined by

$$\bar{\rho}(\rho_L, \rho_R) = \begin{cases} \frac{\rho_R - \rho_L}{\ln(\rho_R) - \ln(\rho_L)} & \text{if } \rho_L \neq \rho_R, \\ \rho_L & \text{if } \rho_L = \rho_R. \end{cases}$$

Then the relaxation scheme preserves exactly the hydrostatic atmosphere:

if the initial data  $w_i^0$  is given by

$$\begin{cases} u_i^0 = 0, \\ \rho_i^0 = \alpha e^{-\beta \Phi_i}, \\ p_i^0 = \frac{\alpha}{\beta} e^{-\beta \Phi_i}, \end{cases}$$

then  $w_i^{n+1} = w_i^n, \quad \forall i \in \mathbb{Z}, \quad \forall n \in \mathbb{N}.$

## Theorem

Assume the parameter  $a$  satisfy the following Whitham conditions:

$$a^2 > p(\tau_{L,R}, e_{L,R}) \partial_e p(\tau_{L,R}, e_{L,R}) - \partial_\tau p(\tau_{L,R}, e_{L,R}),$$

$$a^2 > p(\tau_{L,R}^*, e_{L,R}^*) \partial_e p(\tau_{L,R}^*, e_{L,R}^*) - \partial_\tau p(\tau_{L,R}^*, e_{L,R}^*).$$

Then the relaxation scheme satisfies

$$\rho_i^{n+1} \mathcal{F}(s_i^{n+1}) \leq \rho_i^n \mathcal{F}(s_i^n) - \frac{\Delta t}{\Delta x} \left( \{\rho \mathcal{F}(s) u\}_{i+1/2}^n - \{\rho \mathcal{F}(s) u\}_{i-1/2}^n \right)$$

for all smooth function  $\mathcal{F}$  such that  $w \mapsto \rho \mathcal{F}(s)$  is strictly convex.

The entropy numerical flux is defined by

$$\{\rho \mathcal{F}(s) u\}_{i+1/2}^n = F_{i+1/2}^\rho \times \begin{cases} \mathcal{F}(s_i^n) & \text{if } F_{i+1/2}^\rho > 0, \\ \mathcal{F}(s_{i+1}^n) & \text{if } F_{i+1/2}^\rho < 0, \end{cases}$$

with  $F_{i+1/2}^\rho$  the  $\rho$ -component of the numerical flux.

# Sketch of the proof (1)

- 1 We introduce

$$I(\pi, \tau) = \pi + a^2 \tau \quad \text{and} \quad J(\pi, e) = e - \frac{\pi^2}{2a^2}.$$

$I$  and  $J$  are strong Riemann invariants, i.e they satisfy

$$\partial_t \Psi(I, J) + u \partial_x \Psi(I, J) = 0,$$

for all smooth function  $\Psi : \mathbb{R}^2 \rightarrow \mathbb{R}$ .

- 2 There exists functions  $\overline{\tau}(I, J)$  and  $\overline{e}(I, J)$  which satisfy

$$\overline{\tau} \left( I_{|\pi=p(\tau,e)}, J_{|\pi=p(\tau,e)} \right) = \tau \quad \text{and} \quad \overline{e} \left( I_{|\pi=p(\tau,e)}, J_{|\pi=p(\tau,e)} \right) = e$$

- 3 We define  $\overline{s}(W) = s(\overline{\tau}(I, J), \overline{e}(I, J))$ . This function reaches its minimum on the equilibrium manifold:

$$\overline{s}(W) \geq \overline{s} \left( W_{|\pi=p(\tau,e)} \right) = s \left( W_{|\pi=p(\tau,e)} \right).$$

## Sketch of the proof (2)

- ④  $\bar{s}$  only depends on  $I$  and  $J \Rightarrow \partial_t \rho \mathcal{F}(\bar{s}) + \partial_x \rho \mathcal{F}(\bar{s}) u = 0$ .

We integrate this equation on  $[x_{i-1/2}, x_{i+1/2}] \times [t^n, t^{n+1}]$ :

$$\begin{aligned} \int_{x_{i-1/2}}^{x_{i+1/2}} (\rho \mathcal{F}(\bar{s})) (W_{\Delta x}(x, t^{n+1})) dx &= \int_{x_{i-1/2}}^{x_{i+1/2}} (\rho \mathcal{F}(\bar{s})) (W_{\Delta x}(x, t^n)) dx \\ &\quad - \int_{t^n}^{t^{n+1}} (\rho \mathcal{F}(\bar{s}) u) (W_{\Delta x}(x_{i+1/2}, t)) dt \\ &\quad + \int_{t^n}^{t^{n+1}} (\rho \mathcal{F}(\bar{s}) u) (W_{\Delta x}(x_{i-1/2}, t)) dt. \end{aligned}$$

We deduce

$$\begin{aligned} \frac{1}{\Delta x} \int_{x_{i-1/2}}^{x_{i+1/2}} (\rho \mathcal{F}(\bar{s})) (W_{\Delta x}(x, t^{n+1})) dx &= \rho_i^n \mathcal{F}(s_i^n) \\ &\quad - \frac{\Delta t}{\Delta x} \left( \{\rho \mathcal{F}(s) u\}_{i+1/2}^n - \{\rho \mathcal{F}(s) u\}_{i-1/2}^n \right). \end{aligned}$$

# Sketch of the proof (3)

⑤ By definition of the scheme:

$$\rho_i^{n+1} \mathcal{F}(s_i^{n+1}) = (\rho \mathcal{F}(s)) \left( \frac{1}{\Delta x} \int_{x_{i-1/2}}^{x_{i+1/2}} W_{\Delta x}(x, t^{n+1})|_{\pi=p(\tau,e)} dx \right).$$

Jensen inequality ( $w \mapsto \rho \mathcal{F}(s)$  is convex):

$$\rho_i^{n+1} \mathcal{F}(s_i^{n+1}) \leq \frac{1}{\Delta x} \int_{x_{i-1/2}}^{x_{i+1/2}} (\rho \mathcal{F}(s)) (W_{\Delta x}(x, t^{n+1})|_{\pi=p(\tau,e)}) dx.$$

$\mathcal{F}$  is increasing + the entropy is minimal on the equilibrium manifold:

$$\rho_i^{n+1} \mathcal{F}(s_i^{n+1}) \leq \frac{1}{\Delta x} \int_{x_{i-1/2}}^{x_{i+1/2}} (\rho \mathcal{F}(\bar{s})) (W_{\Delta x}(x, t^{n+1})) dx.$$

# Hydrostatic atmosphere

- Computational domain:  $[0, 1]$  until time  $T = 0.25$

- $\Phi(x) = x^2$

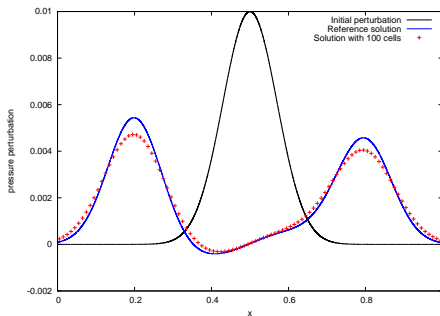
- Initial condition: 
$$\begin{cases} \rho_0(x) = e^{-x^2} \\ u_0(x) = 0 \\ p_0(x) = e^{-x^2} \end{cases}$$

- $L^1$  error:

$N$	Density	Velocity
100	1.15E-16	7.62E-17
200	1.77E-16	1.34E-16
400	3.01E-16	1.14E-16
800	4.37E-16	1.33E-16
1600	7.32E-16	1.91E-16
3200	1.18E-15	2.52E-16

# Small perturbation of an hydrostatic atmosphere

- Computational domain:  $[0, 1]$  until time  $T = 0.25$
- $\Phi(x) = x$
- Initial condition: 
$$\begin{cases} \rho_0(x) = e^{-x} \\ u_0(x) = 0 \\ p_0(x) = e^{-x} + 0.01e^{100(x-0.5)^2} \end{cases}$$
- Pressure perturbation:  $\delta p = p - p_0$





# Non-hydrostatic steady state

- Computational domain:  $[0, 1]$  with periodic boundary conditions until time  $T = 1$
- $\Phi(x) = \sin(2\pi x)$
- Initial condition: 
$$\begin{cases} \rho_0(x) = 3 + 2 \sin(2\pi x) \\ u_0(x) = 0 \\ p_0(x) = 3 + 3 \sin(2\pi x) - \frac{1}{2} \cos(4\pi x) \end{cases}$$

We check easily  $\partial_x p_0 + \rho_0 \partial_x \Phi = 0$

- $L^1$  error:

$N$	Density		Velocity	
100	4.46E-05	—	2.03E-05	—
200	7.11E-06	2.65	5.29E-06	1.94
400	1.23E-06	2.53	1.34E-06	1.98
800	2.35E-07	2.39	3.37E-07	1.99
1600	5.02E-08	2.23	8.44E-08	2.00
3200	1.15E-08	2.13	2.11E-08	2.00

# Perspectives

- Extension to second-order
  - ▶ MUSCL technique written as a convex combination of first-order scheme
  - ▶ Difficulty to preserve the well-balanced property
  - ▶ Possible solution: use a unusual set of variables for the reconstruction
- Extension to unstructured 2D
  - ▶ Convex combination of 1D scheme by interface